Supporting Information

High-Efficiency and Thermal/Moisture Stable CsPbI_{2.84}Br_{0.16} Inorganic Perovskite Solar Cells Enabled by a Multifunctional Cesium Trimethylacetate Organic Additive

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Supporting Figures and Tables



Figure S1. (a) The enlarged view of the (100) peak for a 0-CsTa film. (b) The enlarged view of the (002) and (110) peaks for a 0.1-CsTa film. (c) The simulated XRD patterns for the α and γ phases.



Figure S2. The (004) and (220) peak intensities for 0.1-CsTa, 0.2-CsTa and 0.3-CsTa films. The red line shows the ratio of the (004) peak intensity to the (220) peak intensity for 0.1-CsTa, 0.2-CsTa and 0.3-CsTa films.



Figure S3. The corresponding inverse pole figures for (a) 0-CsTa and (b) 0.1-CsTa films. (c) The color key (inserted in Figure 1d) varies with the orientation of the crystal. The red color in the coordinate represents the growth orientation perpendicular to the substrate. The other colors represent the growth in the other directions.



Figure S4. The UV-Vis spectra for n-CsTa CsPbI_{2.84}Br_{0.16} perovskite films.



Figure S5. The size distribution of the crystals in n-CsTa films determined by counting 50 crystal grains. (a) 0-CsTa. (b) 0.1-CsTa. (c) 0.2-CsTa. (d) 0.3-CsTa.



Figure S6. The binding energies for (a) I 3d, (b) Br 3d and (c) Cs 3d for n-CsTa films tested with XPS. (d) The survey curves for the XPS spectra.



Figure S7. (a) Images for the solubility of PbI₂, CsTa and PbI₂/CsTa mixture powders in DMSO. (b) XRD patterns for films obtained from PbI₂ and PbI₂/CsTa precursor solutions. The black film preparation processes: PbI₂/CsTa precursor solution was spin-coated on a 150°C preheated substrate. Then the black film was obtained by annealing the wet-fresh film at 105°C. (c) The reaction processes and the schematic for γ -phase perovskite formation in PbI₂/CsTa case.



Figure S8. The intensity-sputtering time and intensity-sputtering depth profiles for the O element and $[PbI_3]$ ⁻ in 0.1-CsTa films. Sputtering rate: 0.23 nm/s.



Figure S9. The XRD patterns for 0.1-CsAc, 0.1-CsPro and 0.1-CsTa films. The black line shows the XRD data for the 0-CsTa perovskite film.



Figure S10. The schematics for the blocking effect due to the terminal groups in CsAc, CsPro and CsTa on the stability of the perovskite structure.



Figure S11. The cross-section SEM images for 0-CsTa and 0.1-CsTa CsPbI $_{2.84}$ Br $_{0.16}$ perovskite films.



Figure S12. The steady output current densities at the maximum PCE points for 0.1-CsTa devices based on the use of PbI₂ as the lead source.



Figure S13. The statistical photoelectric parameters for 0.1-CsTa PSCs (40 devices).



Figure S14. Photovoltaic parameter decays for PbI_2 -prepared 0-CsTa PSCs for environmental stability testing in air (RH: ~20%, RT: 10°C).



Figure S15. Photovoltaic parameter decays for 0.1-CsTa and 0.2-CsTa PSCs for environmental stability testing in air (RH: ~20%, RT: 10°C).



Figure S16. Photovoltaic parameter decays for PbI₂-prepared 0-CsTa PSCs for thermal stability testing under a RH=80% and T=80°C environment.



Figure S17. The XRD patterns for 0.1-CsTa and 0.2-CsTa perovskite films obtained under the same environment for stability tests. (a) 0.1-CsTa and (b) 0.2-CsTa perovskite films under a RH= \sim 20% and T=10°C, environment for 1000 hours and 1900 hours. (c) 0.1-CsTa and (d) 0.2-CsTa perovskite films under a RH=80% and T=80°C environment before and after the same storage time of 360 hours.



Figure S18. The statistical photoelectric parameters obtained for 0.1-CsTa PSCs based on the use of HPbI₃ as the lead source from measurements for 20 devices.

Table S1 Data obtained by fitting TRPL spectra for films without ETL based on 0-CsTa and 0.1-CsTa precursor solutions with an exponential decay function.

| Samples | $\tau_1(ns)$ | T ₁ (%) | $\tau_2(ns)$ | $T_2(\%)$ | τ_3 (ns) | T ₃ (%) | $\tau_{avg}\left(ns\right){}^{a}$ |
|---------------------------|---------------------|--------------------|--------------|-----------|---------------|--------------------|-----------------------------------|
| 0-CsTa | 0.97 | 74.91 | 4.55 | 25.09 | 0 | 0 | 1.86 |
| 0.1-CsTa | 1.33 | 9.31 | 27.30 | 25.35 | 156.35 | 65.34 | 109.20 |
| $\tau_{avg} = \sum_{i} I$ | $A_i \tau_i$, wher | $e\sum_{i}A_{i}=1$ | | | | | |

Table S2 Photovoltaic parameters for PSCs based on n-CsTa PbI₂-based precursors

| Samples | Scanning | J_{sc} | V_{oc} | FF | PCE | Hysteresis |
|-------------------|------------|-----------------------|----------|-------|-------|------------|
| n-CsTa | directions | [mA/cm ²] | [V] | [%] | [%] | [%] |
| 0 C-T- | R | 13.18 | 0.85 | 67.63 | 7.61 | 41.65 |
| 0-Cs1a | F | 14.51 | 0.83 | 36.84 | 4.44 | - |
| | R | 17.45 | 1.09 | 77.28 | 14.70 | 15.3 |
| 0.1 - CSTa | F | 17.36 | 1.09 | 65.68 | 12.45 | - |
| | R | 15.68 | 1.09 | 69.06 | 11.85 | 17.0 |
| 0.2 - CSTa | F | 16.15 | 1.02 | 59.58 | 9.84 | |
| | R | 14.87 | 0.91 | 68.14 | 9.23 | 13.4 |
| 0.3-0818 | F | 14.65 | 0.83 | 65.04 | 7.94 | - |

Table S3 Photovoltaic parameters of PSCs based on HPbI₃-prepared 0-CsTa and 0.1-CsTa precursor solutions.

| Samples | Scanning | $J_{\scriptscriptstyle SC}$ | V _{oc} | FF | PCE |
|----------|------------|-----------------------------|-----------------|-------|-------|
| 0.1-CsTa | directions | [mA/cm ²] | [V] | [%] | [%] |
| 0-CsTa | R | 19.23 | 1.05 | 75.43 | 15.25 |

| | F | 17.44 | 0.98 | 75.13 | 12.93 |
|----------|---|-------|------------|-------|-------|
| | R | 19.48 | 1.10 77.30 | 16.59 | |
| 0.1-CSTa | F | 18.29 | 1.07 | 71.93 | 14.09 |

Table S4 Stability results obtained from representative studies for organic-added CsPbI₃ PSCs without encapsulation

| Perovskite/Organic additives | PCE [%] | Storage environments and times | Remaining the initial PCE | Reference |
|---------------------------------|------------|--------------------------------------|---------------------------|--|
| CsPbI ₃ /zwitterion | 11.4 | RH: 30% 720 hours | 85% | <i>Joule</i> , 2017 , 1(2): 371-382. |
| CsPbI3/BAI | 4.8 | RH: 30% 720 hours | 92% | Journal of Materials Chemistry A, 2017 , 5(5): 2066- 2072. |
| CsPbI ₃ /EDAI | 11.8 | Dark dry air 720 hours | 84.3% | <i>Science</i> <i>advances</i> , 2017 , 3(9): e1700841. |
| CsPbI ₃ /PEAI | 5.7 | Dry air 96 hours | 75% | Nanoscale, 2018, 10(14): 6318-6322. |
| CsPbI ₃ /PEAI | 12.4 | RH:20±5% 960 hours | 83% | <i>Joule</i> , 2018 , 2(7): 1356-1368. |
| CsPbI ₃ / PEAI | 15.1 | RH: 20-30% 1440 hours | 92% | Nature communica |

| | | | | tions, | |
|--|-------|--|-----------------|--|--|
| | | | | 2018 , 9(1): | |
| | | | | 1-8. | |
| CsPbI ₃ /PVP | 10.7 | RH: 45-55% 500 hours | 75% | Nature communica tions, 2018, 9(1): 1-8. | |
| CsPbI3/DETAI | 7.9 | Dark dry air 1008 hours | 92% | Journal of Materials Chemistry A, 2018 , 6(37): 18258- 18266. | |
| CsPbI ₃ /OTG3 | 13.2 | Dark RH: ~10% 720 hours | 85% | Advanced materials, 2019 , 31(24): 1900605. | |
| CsPbI ₃ /MAI/MAS | 16.04 | Nitrogen 1000 hours | 95% | <i>ACS</i> <i>Energy</i> <i>Letters</i> , 2020 , 5(5): 1619-1627. | |
| CsPbI _{2.84} Br _{0.16} /0.1-CsTa (lead source: PbI ₂) | 14.70 | RH: 80% Temperature:80° C 360 hours | 64.07% | This work | |
| | | RH: 20% 1008 hours/ 1968 hours | 88.5%/ 79.8% | | |
| CsPbI _{2.84} Br _{0.16} /0.2-CsTa | 11.85 | RH: 80% | 87.4% | | |

| | | Temperature:80° | | |
|--|-------|-----------------|-----------|--|
| | | С | | |
| (load source: DhI) | | 500 hours | | |
| $(1ead source. For_2)$ | | RH: 20% | 05 269/ / | |
| | | 1008 hours / | 93.20% | |
| | | 1968 hours | 89.0870 | |
| CsPbI _{2.84} Br _{0.16} /0.1-CsTa | 16.50 | RH: 20% | 20.020/ | |
| (lead source: HPbI ₃) | | 1296 hours | 80.88% | |
| | | | | |

Experimental sections

1. Precursor Solution Preparation

The CsTa added CsPbI_{2.84}Br_{0.16} perovskite precursor solutions were prepared by using mole ratios of CsI: PbI₂: CsBr: CsTa= 1:1:0.4:n (n=0, 0.1, 0.2 and 0.3). m mol (m=0, 0.08, 0.16 and 0.24) CsTa, 0.8 mol cesium iodide (CsI, 99%) and 0.32 mol cesium bromide (CsBr, 99.99%) were dissolved with a 0.8 mol lead source into 1 ml DMSO to yield 0.8 M precursor solutions. The precursor solutions were stirred at 60 °C for 4 hours, and then filtered with a 0.25 μ m nylon filter. The raw materials, PbI₂, CsBr, CsI and CsTa, were purchased from TCI. All the materials were used as received without further purification. HPbI₃ was synthesized by adding PbI₂ and 57 wt% hydroiodic acid with a molar ratio of 1:1.4 into DMF. The mixed solution was stirred for 2 h. The precipitate was then obtained by evaporating the solutions at 70 °C. The resulting solid was washed four times with copious diethyl ether until the supernatant turned colorless. Finally, the powders were dried in a vacuum oven at 50 °C.

2. Device Fabrication

The FTO was cleaned with detergent, deionized water, and ethanol. The cleaned substrate was then treated with UV ozone for 20 min. The compact TiO_2 (c- TiO_2) was fabricated by dipping the FTO into a 200 ml cold aqueous solution of 4.5 ml $TiCl_4$ at 70°C for 1 h. After that, the substrate was annealed at 200°C for 30 mins. The FTO/c-

TiO₂ was then treated with UV ozone for 20 min once more. Then, 45 µl perovskite precursor solutions were drop cast onto FTO/c-TiO₂ preheated to 90°C. Next, the wet films were immediately spun at 3000 rpm for 30 s, and then annealed at 60 °C for 3 min and 180 °C for 3 min. For the HTL, a solution of 72.3 mg spiro-MeOTAD, 17.5 mL of a prepared solution of 520 mg/mL lithium bis(trifluoromethylsulphonyl)imide, and 28.8 mL 4-tert-butylpyridine in 1 mL chlorobenzene was spun at 3000 rpm for 20 s. The devices were completed by depositing a 75 nm Au cathode by thermal evaporation under a pressure of $4 \times 10-4$ Pa. Both PbI₂ and HPbI₃-based devices were prepared by following the above fabrication processes except for the use of an annealing time of 20 min at 180°C for the HPbI₃-based devices.

3. Device Characterization

XRD patterns were obtained by using an X-ray diffractometer (SmartLab, Rigaku) with Cu Ka radiation (1.5418 Å). SEM images were taken with a SU8010 SEM (Hitachi). Absorption spectra were measured with a UV-2450 spectrophotometer (Shimadzu). The steady-state fluorescence and TRPL decay spectra were measured with a steady/transient state spectrophotometer (NanoLOG-TCSPC, USA) with an excitation wavelength of 450 nm. FTO with a sheet resistance of 15 Ω per square was obtained from Pilkington TEC. The active area of the solar cell was 0.09 cm², which was defined to be 0.09 cm² by using a black aperture mask. The J-V curves were measured using a Keithley 2400 source meter together with a sunlight simulator (XES-300T1, SANEI Electric, AM 1.5 G), which was calibrated using a standard silicon reference cell (scan rate: 44 mV/s). The incident photon-to-electron conversion efficiency (IPCE) was measured in air using a QE-R measurement system (Enli Technology). The IR spectra were obtained with a Fourier-transform infrared spectrometer (Nicolet iS50, USA). Transient photovoltage decay measurements were performed using an electrochemical workstation (Zahner). X-ray photoelectron spectroscopy (XPS) was performed using an RBD upgraded PHI-5000C ESCA system (Perkin Elmer) with Mg K α radiation (h μ = 1486.6 eV). The contact angles were collected by using a Dataphysics OCA15EC. EBSD tests were measured with a highperformance field emission scanning electron microscope (Zeiss Merlin). TOF-SIMS was performed using a TOF-SIMS5 (ION-TOF GmbH, Germany) by utilizing a cesium-ion beam (Sput. Rate=0.23 nm/s).

4. DFT Calculation Methods

In this work, the density functional theory (DFT) calculation was performed using the Dmol³ code.¹ The exchange-correlation interaction was treated by the generalized gradient approximation (GGA) with the PBE functional.² A double numerical quality basis set with d-type polarization function (DNP³) was utilized for all the geometric optimizations, total energy calculations. The core electrons were modeled using effective core pseudopotentials (ECP) by Dolg⁴ and Bergner⁵. All calculations were spin unrestricted. The positions of all the atoms were fully relaxed until the following convergence criterion were met: 0.002 Ha/Å for force, 10⁻⁵ Ha for total energy and 0.005 Å for displacement. The real space cutoff radius for complete vacancy formation was 5.5 Å. The self-consistent field computation criterion was chosen to be 10⁻⁶ Ha.

5. Space Charge Limited Current Methods

The trap density can be calculated from the following equation

$$V_{TFL} = \frac{en_t L^2}{2\varepsilon\varepsilon_0}$$

where L is the thickness of the perovskite film (\approx 480 nm), ε is the relative dielectric constant of the inorganic perovskite (taken as 18 from a previous report⁶), ε_0 is the vacuum permittivity (8.854×10⁻¹² F/m), and *e* is the unit charge with a value of 1.602×10⁻¹⁹ C. n_t is the trap density.

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