Supplementary Material

## Trimming defective perovskite layer surface for high-

## performance solar cells

Chanhyeok Kim<sup>1</sup>, Kihoon Kim<sup>1</sup>, Youngmin Kim<sup>1</sup>, Nikolai Tsvetkov<sup>1</sup>, Nam Joong Jeon<sup>2</sup>, Bong Joo Kang<sup>2,\*</sup>, Hanul Min<sup>1,3,4,\*</sup>

<sup>1</sup>KU-KIST Graduate School of Converging Science and Technology, Korea University, Seoul, South Korea

<sup>2</sup>Division of Advanced Materials, Korea Research Institute of Chemical Technology, Daejeon, South Korea

<sup>3</sup>Department of Integrative Energy Engineering, Korea University, Seoul, South Korea

<sup>4</sup>Clean Energy Research Center, Korea Institute of Science and Technology, Seoul, South Korea

\*To whom correspondence should be addressed. E-mail: bjkang@krict.re.kr; hmin92@korea.ac.kr



Fig. S1. Cross-sectional SEM images of the control, TS (05:95), TS (10:90), and TS (20:80) devices.



**Fig. S2. (a)** Absorption coefficient of the perovskite layer and **(b)** transmitted light intensity for different wavelengths.



Fig. S3. Atomic force microscopy images of the control, TS (05:95), TS (10:90), and TS (20:80) perovskite thin films.



Fig. S4. Williamson-Hall plots of the control, TS (05:95), TS (10:90), and TS (20:80) perovskite thin films.



Fig. S5. Cross-sectional SEM images of the perovskite film for different perovskite precursor concentrations.



Fig. S6. XRD patterns and individual Williamson-Hall plot of the perovskite thin films prepared with 1.40 M, 1.12 M, 1.05 M, 0.92 M, and 0.70 M of FAPbI<sub>3</sub>.

![](_page_5_Figure_0.jpeg)

Fig. S7. GIXRD patterns of the control, TS (05:95), TS (10:90), and TS (20:80) perovskite thin films at different incidence angles.

![](_page_6_Figure_0.jpeg)

Fig. S8. Magnified GIXRD peaks corresponding to the (100) plane of the control, TS (05:95), TS (10:90), and TS (20:80) perovskite films at varying incidence angles.

![](_page_7_Figure_0.jpeg)

Fig. S9. X-ray penetration depth  $\Lambda$  into the FAPbI<sub>3</sub> layer at different incidence angles,

$$\Lambda = \frac{\lambda}{\sqrt{2}\pi} \times \frac{1}{\sqrt{\sqrt{\left(\alpha_i^2 - \alpha_c^2\right)^2 + 4\beta^2} - \left(\alpha_i^2 - \alpha_c^2\right)}}$$

calculated as:

where  $\alpha_i$  is the incidence angle,  $\alpha_c$  is the critical angle of the sample,  $\lambda$  is the X-ray wavelength, and  $\beta$  is the imaginary part of the refraction index, defined as  $n = 1 - \delta + i\beta$ . The  $\delta$  and  $\beta$  parameters for FAPbI<sub>3</sub> at 8.04 keV ( $\lambda = 1.54$  Å) were 1.08329 × 10<sup>-5</sup> and 1.22906 × 10<sup>-6</sup>, respectively, as determined using the online toolbox available at *https://henke.lbl.gov/optical\_constants/getdb2.html*.

**Table S1.** Microstrains calculated from the GIXRD patterns of the control, TS (05:95), TS (10:90), and TS (20:80) thin films at different incidence angles.

Incidence angle	0.25° (ε x 10 <sup>4</sup> )	0.5° (ε x 10 <sup>4</sup> )	0.75° (ε x 10 <sup>4</sup> )	1° (ε x 10 <sup>4</sup> )	1.5° (ε x 10 <sup>4</sup> )
Control	4.3223	3.8565	1.9579	0.3856	0.2551
TS (05:95)	3.9907	2.9453	1.7198	0.3545	0.2537
TS (10:90)	3.2813	2.8302	1.4861	0.3249	0.2586
TS (20:80)	2.7364	2.1199	1.2695	0.3105	0.2547

![](_page_9_Figure_0.jpeg)

**Fig. S10. (a)** Perovskite film peeled-off from the substrate. **(b-c)** Magnified GIXRD diffraction peaks corresponding to the (100) plane of the buried interface. **(d-g)** WH plots of the buried interface of the control and TS (05:95).

**Table S2.** Microstrains calculated from the GIXRD patterns of the buried interface of the control and TS (05:95) thin films at different incidence angles.

Incidence angle	0.1° (ε x 10 <sup>4</sup> )	0.2° (ε x 10 <sup>4</sup> )	
Control	0.9285	1.3343	
TS (05:95)	0.9099	1.3802	

![](_page_10_Figure_0.jpeg)

Fig. S11. Statistics of key performance parameters of the control, TS (05:95), TS (10:90), and TS (20:80) devices.

![](_page_10_Figure_2.jpeg)

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Fig. S12. Independent PCE certification issued by the Korea Institute of Energy Research.

![](_page_11_Figure_0.jpeg)

**Fig. S13.** Stabilized power output and current density at the maximum power point for the control and TS (05:95) devices.

![](_page_12_Figure_0.jpeg)

Fig. S14. UV-vis absorption spectra and bandgaps of the control, TS (05:95), TS (10:90), and TS (20:80) perovskite thin films.

![](_page_13_Figure_0.jpeg)

Fig. S15. Semilog EQE vs photon energy for the control and TS (05:95) devices.